

PSR J1023+0038

spinning down faster or slower than as a radio pulsar?

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Task force!

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Slavko Bogdanov

Alessandro Patrino

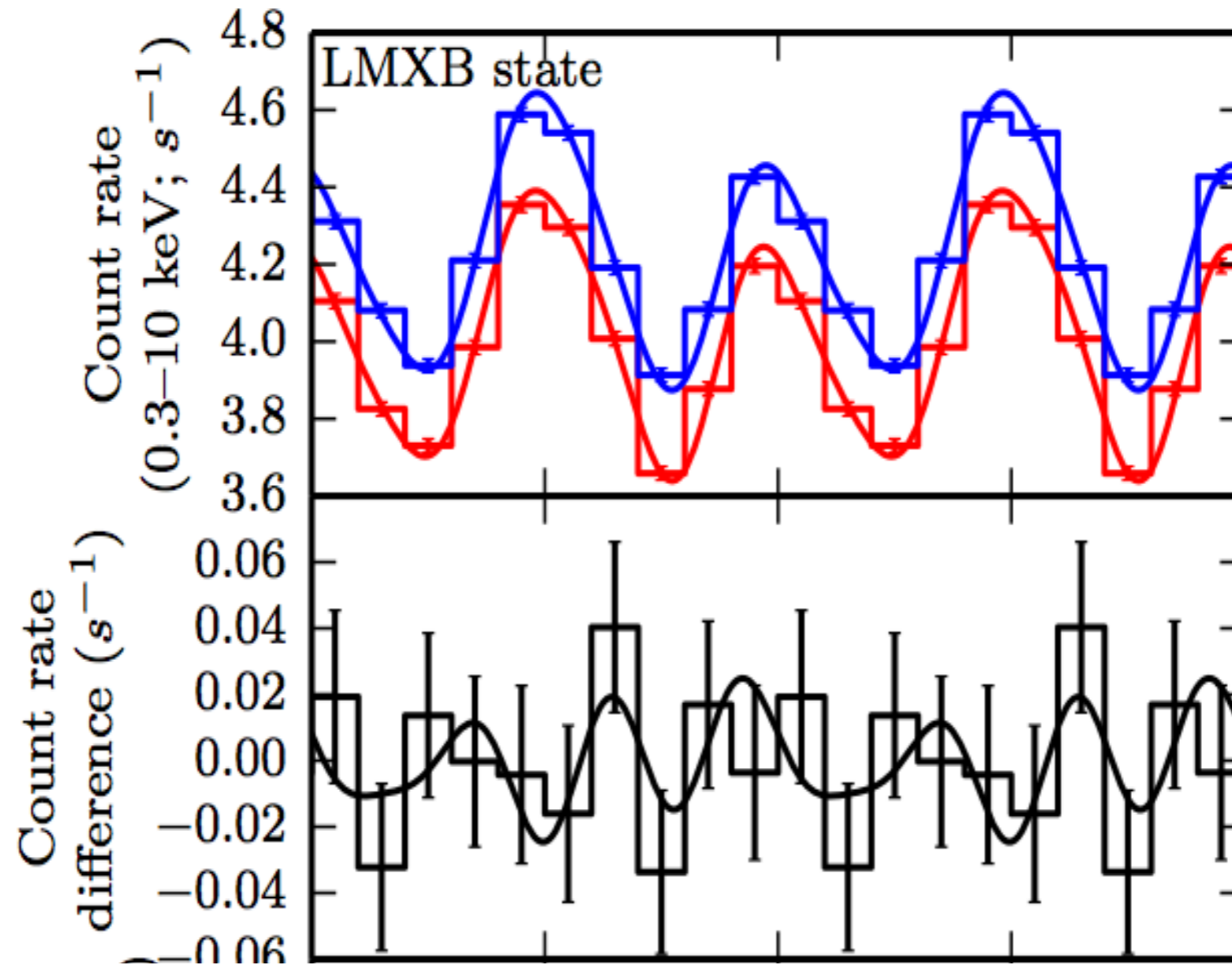
Caroline D'Angelo

Radio Timing Solution

Fit and data-set	
Pulsar name	J1023+0038
MJD range.....	54766.5—56146.6
Number of TOAs.....	7478
Rms timing residual (μs).....	114.0
Weighted fit.....	No
Measured Quantities	
Pulse frequency, ν (s^{-1}).....	592.42145906986(10)
First derivative of pulse frequency, $\dot{\nu}$ (s^{-2})	$-2.432(3) \times 10^{-15}$
Orbital period, P_b (d)	0.1980963569(3)
Epoch of periastron, T_0 (MJD)	54905.9713992(3)
Projected semi-major axis of orbit, x (lt-s)	0.343343(3)
First derivative of orbital period, \dot{P}_b	$-7.32(6) \times 10^{-11}$
Set Quantities	
Right ascension, α	10:23:47.687198
Declination, δ	+00:38:40.84551
Epoch of frequency determination (MJD).....	54906
Epoch of position determination (MJD)	54995
Dispersion measure, DM ($cm^{-3}pc$)	14.3308
Proper motion in right ascension, μ_α ($mas\ yr^{-1}$)	4.76
Proper motion in declination, μ_δ ($mas\ yr^{-1}$).....	-17.34
Parallax, π (mas)	0.000731
Orbital eccentricity, e	0

Archibald et al. 2013, arXiv:1311.5161

Coherent X-ray pulsations



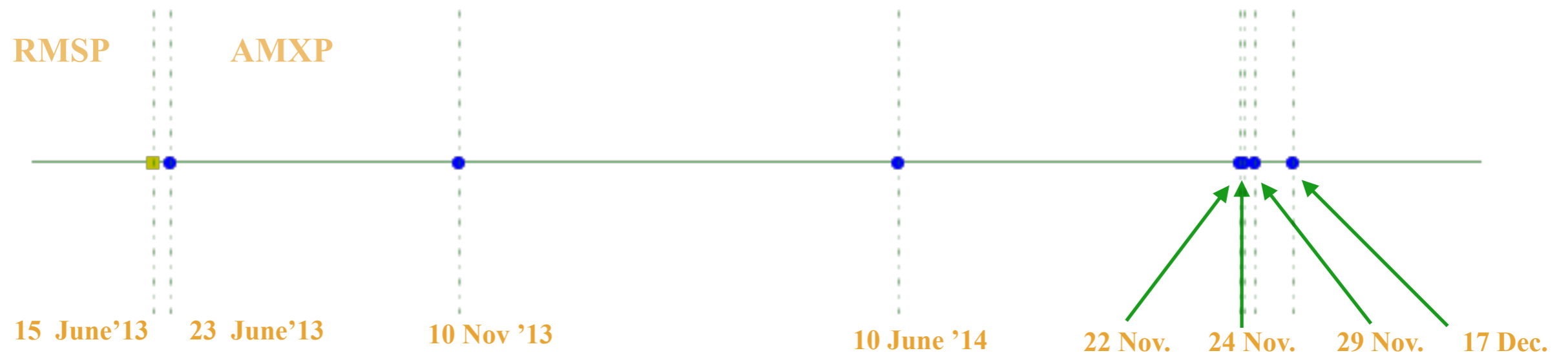
Pulse profiles

Archibald et al. 2014, arXiv:1412.1306
Bogdanov et al. 2014, arXiv:1412.5145

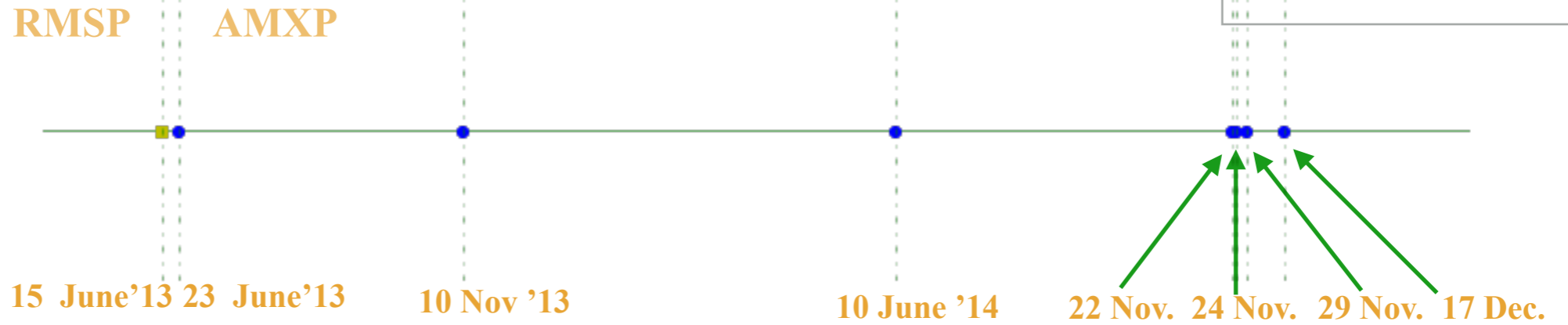
XMM-Newton



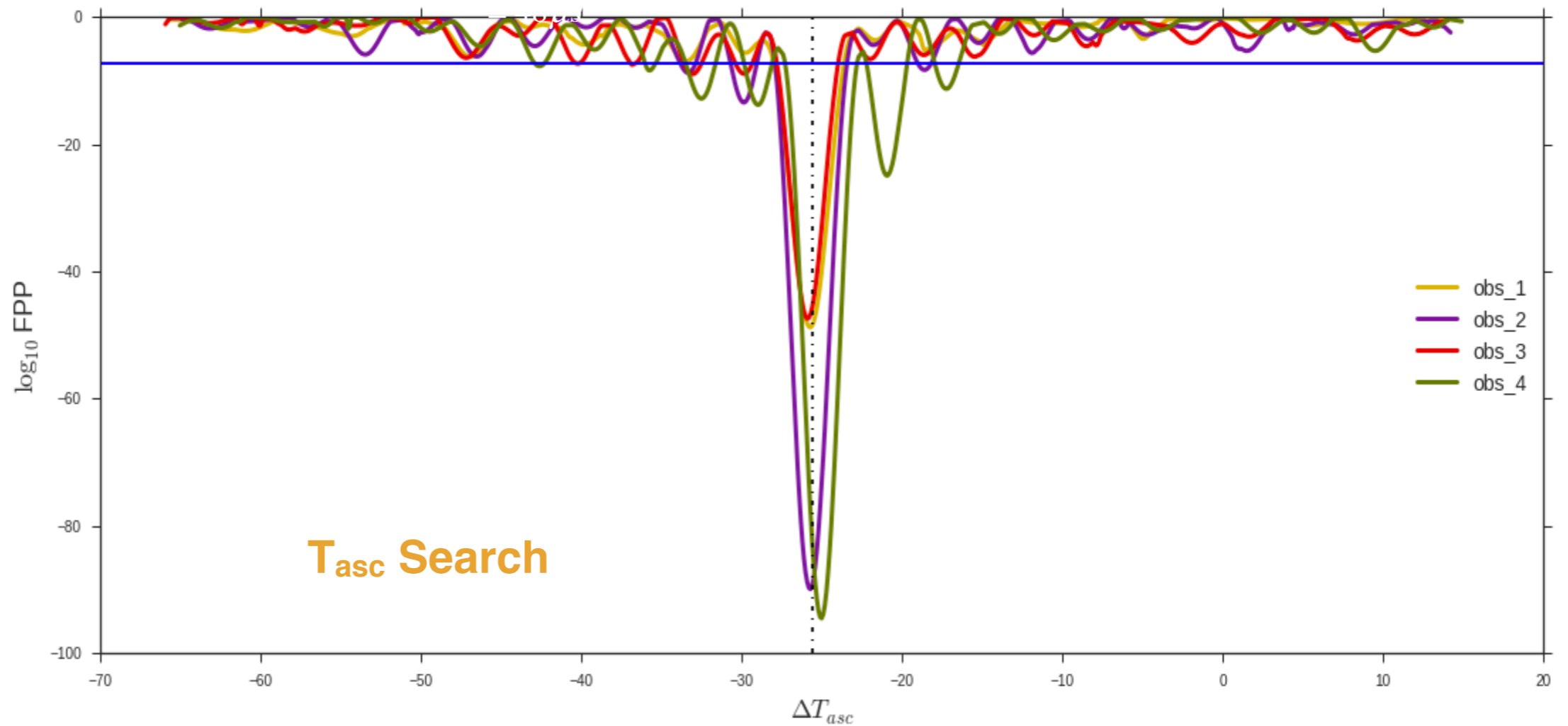
- X-ray Multi-mirror satellite
1) EPIC 2) RGS 3) OM
- Highly eccentric , 48 hour orbit
- 30 arcmin FOV, 6 arcsec resolution, 0.5-15 KeV



T_{asc} search



Pulse profile

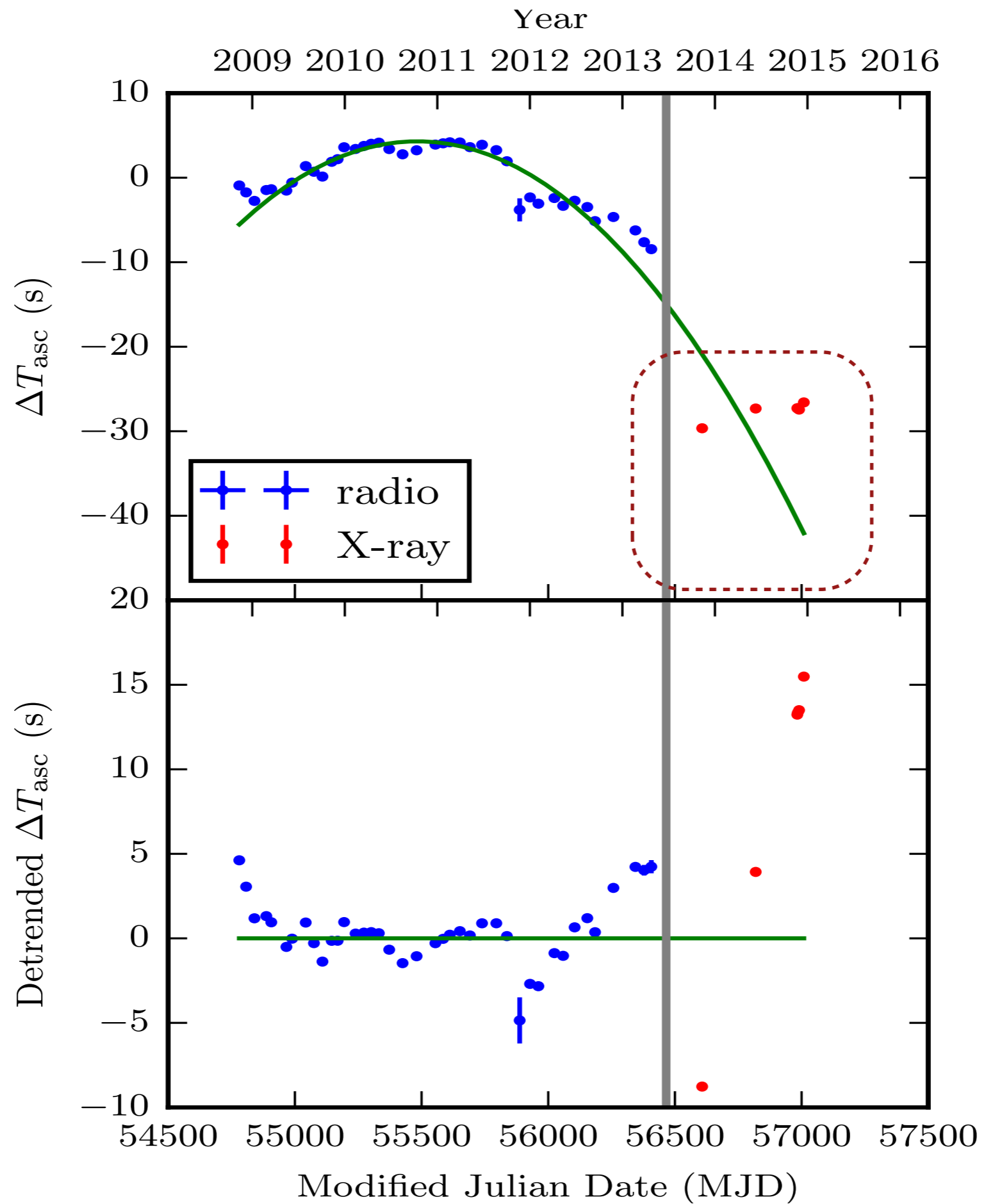


Change in frequency

T_{asc} Search

Period Derivative

Unambiguous presence of pulsations : FPP 10^{-90}

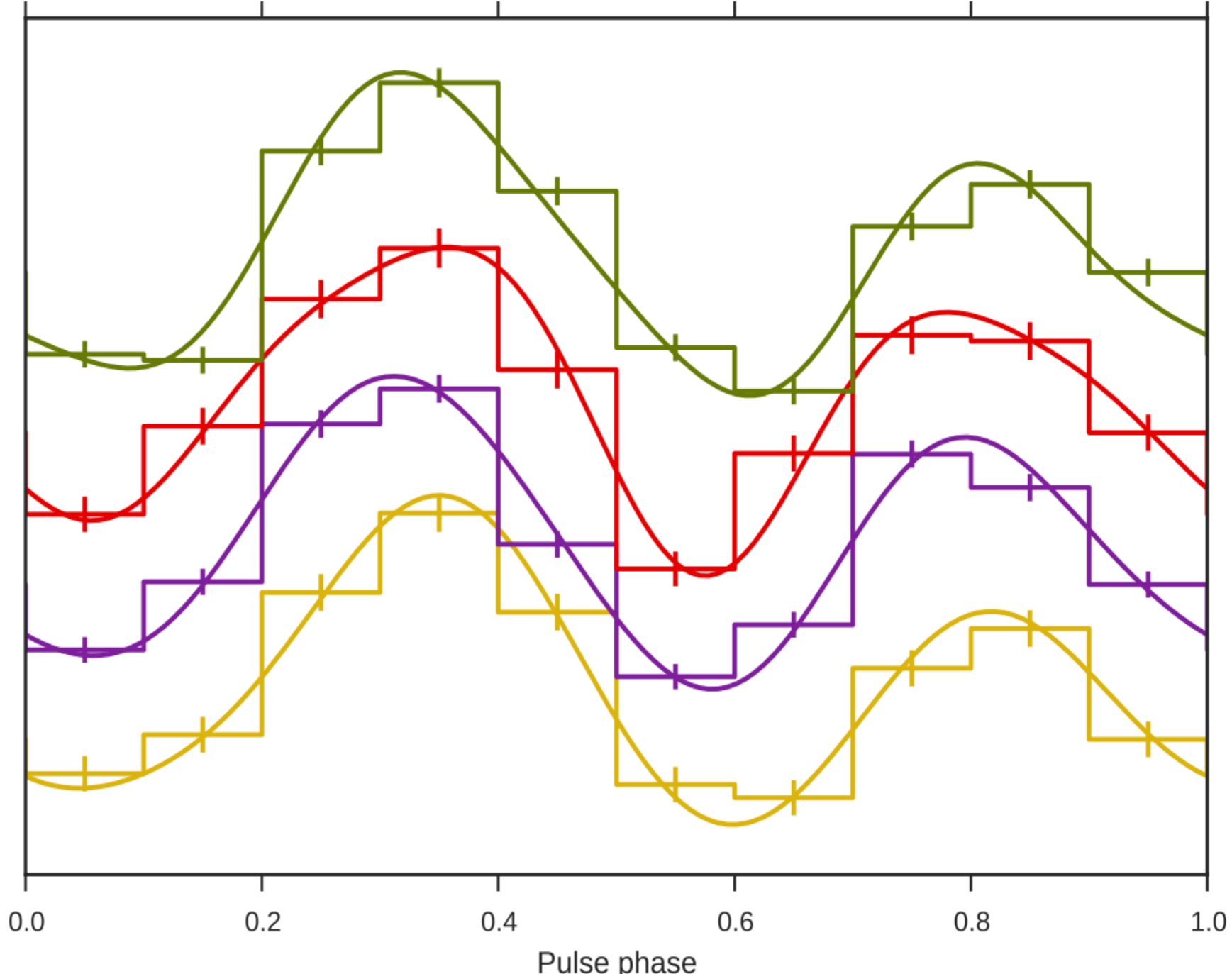


Background subtracted, normalised folded profiles based on the radio ephemerides.

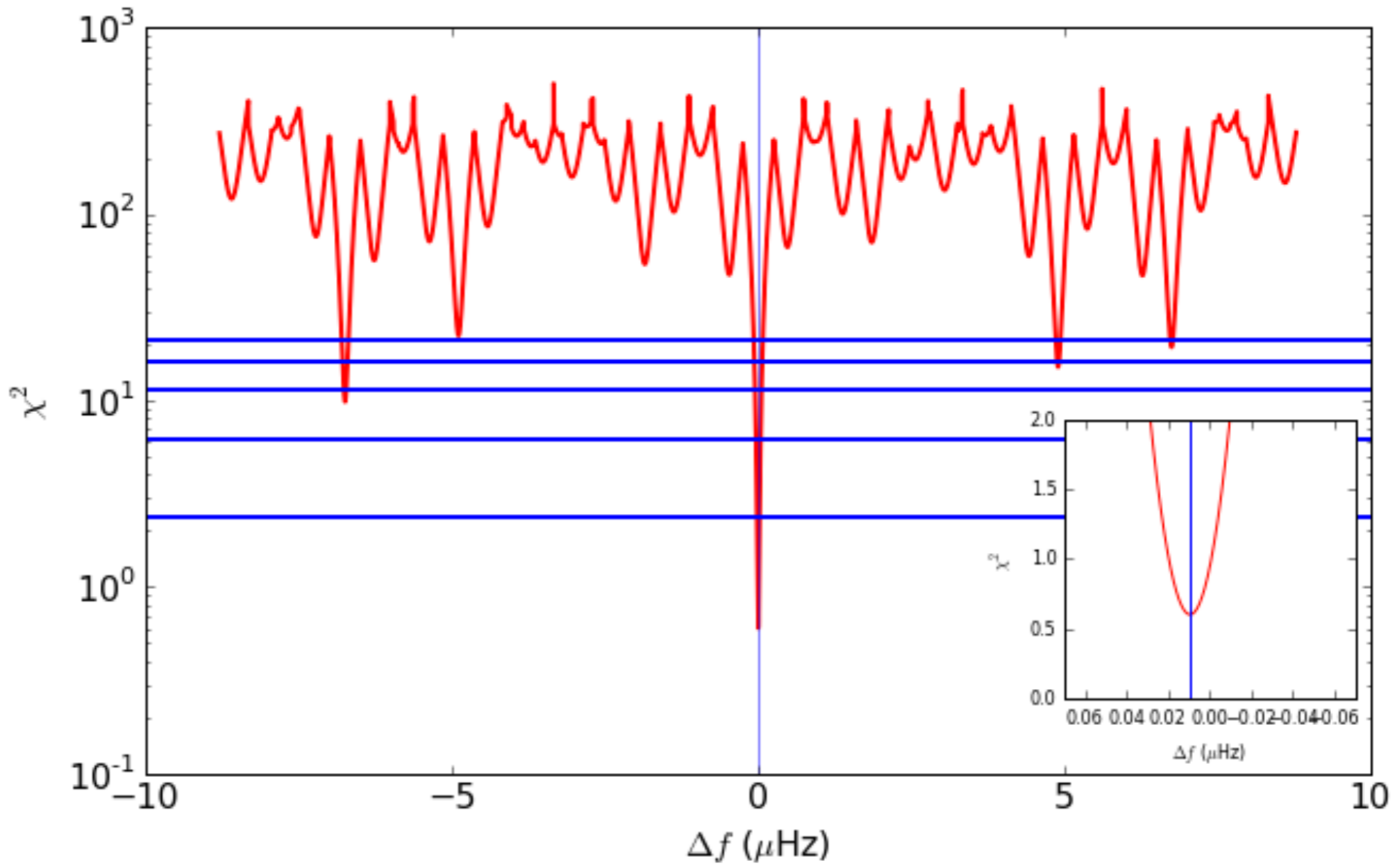
Pulse profile

Change in frequency

Period Derivative



Distribution over change in frequency

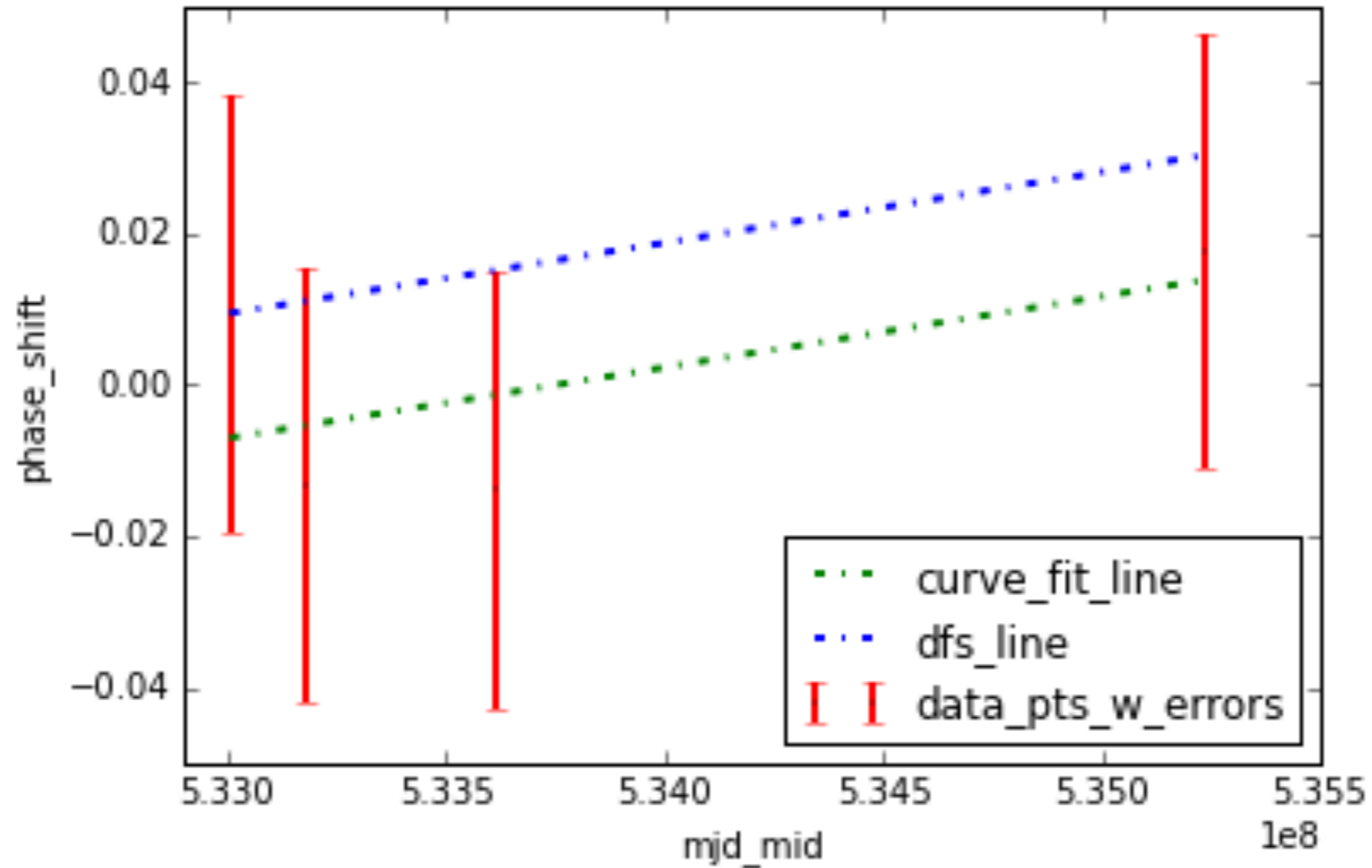


Preliminary key results

T_{asc} search

Pulse profile

Change in frequency

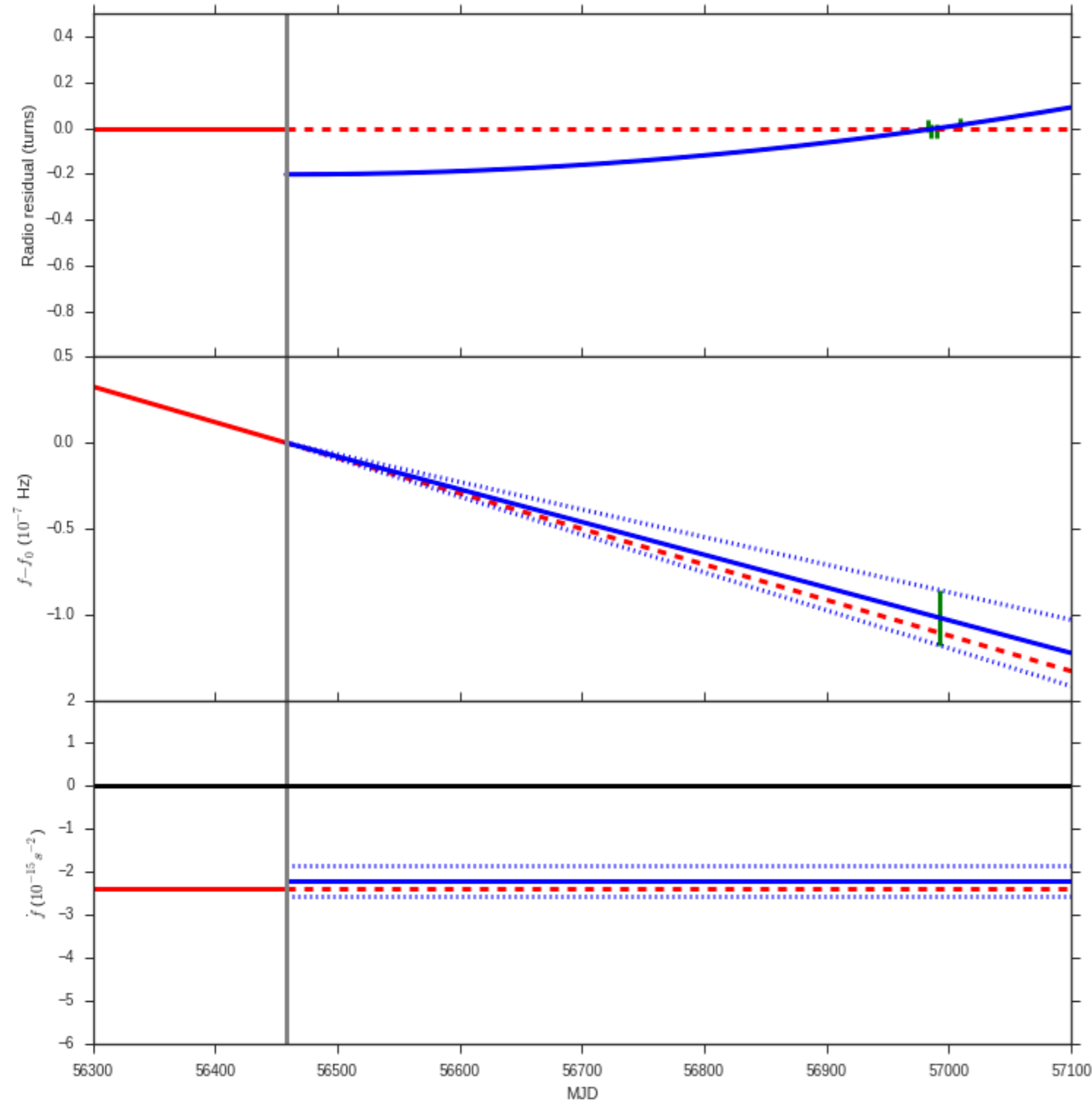


$$\dot{P}_{accr} = \dot{P}_{radio} (1 \pm 0.08)$$

Period
Derivative

Conclusions

$$\dot{f}_{\text{accr}} \approx \dot{f}_{\text{radio}}$$



- If the pulsar spin down mechanism remains active?
- At present is the pulsar wind being generated?
- Dying down of torques
- Is it “mode switching”?
- Possible models?
- See talk by Caroline D’Angelo